

SPHERE data centering using satellite spots

Jules Dallant

Centre de Recherche Astrophysique de Lyon

October 4, 2022



- ① Satellite spots
- ② Estimation of the rotation center
- ③ Application to SPHERE data
- ④ Conclusions & perspectives

Table of Contents

- ① Satellite spots
- ② Estimation of the rotation center
- ③ Application to SPHERE data
- ④ Conclusions & perspectives

In SPHERE's pupil-stabilized mode, the field rotates !

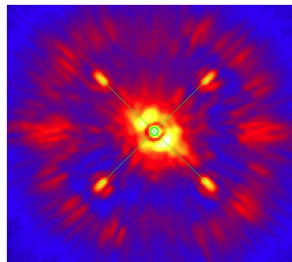
→ exact position of the rotation center is critical for astrometric measurements

Satellite spots

In SPHERE's pupil-stabilized mode, the field rotates !

→ exact position of the rotation center is critical for astrometric measurements

Applying a waffle pattern to the DM creates *satellite spots* :



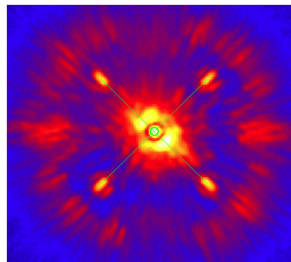
Satellite spots

In SPHERE's pupil-stabilized mode, the field rotates !

→ exact position of the rotation center is critical for astrometric measurements

Applying a waffle pattern to the DM creates *satellite spots* :

- 4 symmetric replicas of the PSF at $\simeq 14\lambda/D$
- displayed either as an \times or $+$ shape
- allows computing the rotation center at the middle of the square pattern
- either elongated (BB) & point sources (NB)



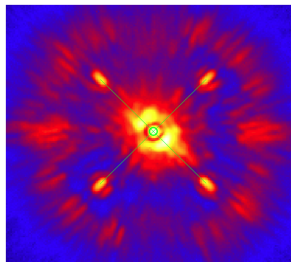
Satellite spots

In SPHERE's pupil-stabilized mode, the field rotates !

→ exact position of the rotation center is critical for astrometric measurements

Applying a waffle pattern to the DM creates *satellite spots* :

- 4 symmetric replicas of the PSF at $\simeq 14\lambda/D$
- displayed either as an \times or $+$ shape
- allows computing the rotation center at the middle of the square pattern
- either elongated (BB) & point sources (NB)



→ Need to **fit** the positions of the **satellite spots** to estimate the rotation center

Table of Contents

- ① Satellite spots
- ② Estimation of the rotation center
- ③ Application to SPHERE data
- ④ Conclusions & perspectives

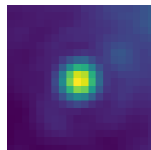
The model

At a given frame k and spectral channel ℓ , satellite spots are considered separately.

The model

At a given frame k and spectral channel ℓ , satellite spots are considered separately.

Data $\mathbf{d} = \{d_i\}_{i=1:N}$, 2-D image of size N :

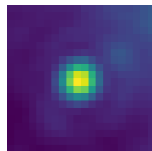


The model

At a given frame k and spectral channel ℓ , satellite spots are considered separately.

Data $\mathbf{d} = \{d_i\}_{i=1:N}$, 2-D image of size N :

2-D Gaussian model $\mathbf{g}(\boldsymbol{\xi}) = \{g_i(\boldsymbol{\xi})\}_{i=1:N}$ of param. $\boldsymbol{\xi}$

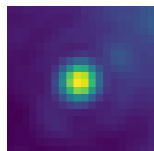


The model

At a given frame k and spectral channel ℓ , satellite spots are considered separately.

Data $\mathbf{d} = \{d_i\}_{i=1:N}$, 2-D image of size N :

2-D Gaussian model $\mathbf{g}(\boldsymbol{\xi}) = \{g_i(\boldsymbol{\xi})\}_{i=1:N}$ of param. $\boldsymbol{\xi}$



Goal :

Solve the following non-linear least mean square constrained optimization problem:

$$\hat{\boldsymbol{\xi}} = \min_{\boldsymbol{\xi}} \left\{ \sum_{i=1}^N (d_i - g_i(\boldsymbol{\xi}))^2 : \boldsymbol{\xi}_l \leq \boldsymbol{\xi} \leq \boldsymbol{\xi}_u \right\}$$

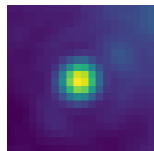
To get the best **estimator for the center coordinates** (x_0, y_0) of the satellite spot.

The model

At a given frame k and spectral channel ℓ , satellite spots are considered separately.

Data $\mathbf{d} = \{d_i\}_{i=1:N}$, 2-D image of size N :

2-D Gaussian model $\mathbf{g}(\boldsymbol{\xi}) = \{g_i(\boldsymbol{\xi})\}_{i=1:N}$ of param. $\boldsymbol{\xi}$



Goal :

Solve the following non-linear least mean square constrained optimization problem:

$$\hat{\boldsymbol{\xi}} = \min_{\boldsymbol{\xi}} \left\{ \sum_{i=1}^N (d_i - g_i(\boldsymbol{\xi}))^2 : \boldsymbol{\xi}_l \leq \boldsymbol{\xi} \leq \boldsymbol{\xi}_u \right\}$$

To get the best **estimator for the center coordinates** (x_0, y_0) of the satellite spot.

Solver: STIR algorithm (Subspace Trust region Interior Reflective)

General 2-D elliptical Gaussian function

The general 2-D elliptical Gaussian function of parameter ξ :

General 2-D elliptical Gaussian function

The general 2-D elliptical Gaussian function of parameter ξ :

$$g(x, y, \xi) = A \exp \left(- \left(\frac{\cos^2 \theta}{2\sigma_x^2} + \frac{\sin^2 \theta}{2\sigma_y^2} \right) (x - x_0)^2 - \left(\frac{\sin^2 \theta}{2\sigma_x^2} + \frac{\cos^2 \theta}{2\sigma_y^2} \right) (y - y_0)^2 - 2 \left(\frac{\sin 2\theta}{4\sigma_x^2} - \frac{\sin 2\theta}{4\sigma_y^2} \right) (x - x_0)(y - y_0) \right) + c$$

General 2-D elliptical Gaussian function

The general 2-D elliptical Gaussian function of parameter ξ :

$$g(x, y, \xi) = A \exp \left(- \left(\frac{\cos^2 \theta}{2\sigma_x^2} + \frac{\sin^2 \theta}{2\sigma_y^2} \right) (x - x_0)^2 - \left(\frac{\sin^2 \theta}{2\sigma_x^2} + \frac{\cos^2 \theta}{2\sigma_y^2} \right) (y - y_0)^2 - 2 \left(\frac{\sin 2\theta}{4\sigma_x^2} - \frac{\sin 2\theta}{4\sigma_y^2} \right) (x - x_0)(y - y_0) \right) + c$$

where $\xi = [A, x_0, y_0, \sigma_x, \sigma_y, \theta, c]^T$ with

- A , the amplitude
- x_0, y_0 , the center coordinates
- σ_x, σ_y , the x and y spreads of the blob
- θ , the rotation angle
- c , an offset (constant)

General 2-D elliptical Gaussian function

The general 2-D elliptical Gaussian function of parameter ξ :

$$g(x, y, \xi) = A \exp \left(- \left(\frac{\cos^2 \theta}{2\sigma_x^2} + \frac{\sin^2 \theta}{2\sigma_y^2} \right) (x - x_0)^2 - \left(\frac{\sin^2 \theta}{2\sigma_x^2} + \frac{\cos^2 \theta}{2\sigma_y^2} \right) (y - y_0)^2 - 2 \left(\frac{\sin 2\theta}{4\sigma_x^2} - \frac{\sin 2\theta}{4\sigma_y^2} \right) (x - x_0)(y - y_0) \right) + c$$

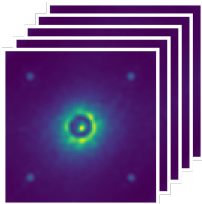
where $\xi = [A, x_0, y_0, \sigma_x, \sigma_y, \theta, c]^T$ with

- A , the amplitude
- x_0, y_0 , the center coordinates
- σ_x, σ_y , the x and y spreads of the blob
- θ , the rotation angle
- c , an offset (constant)

The discrete 2-D elliptical Gaussian function at pixel i is:

$$g_i(\xi) = g(x_i, y_i, \xi)$$

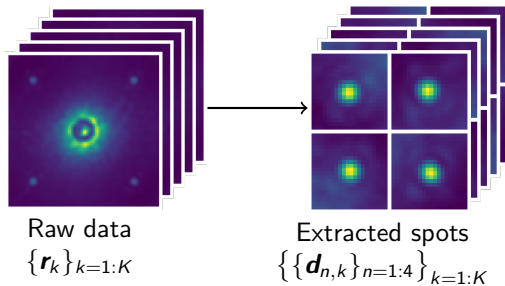
Extracting & fitting the satellite spots



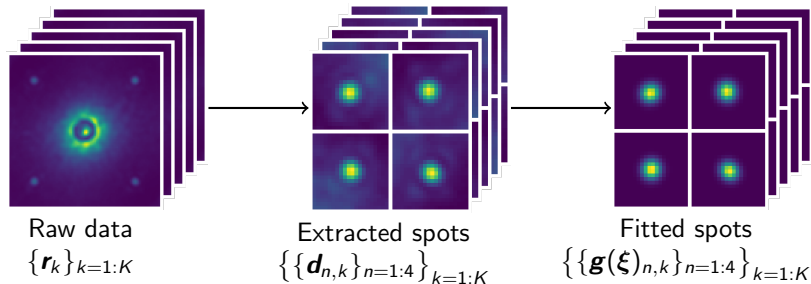
Raw data

$$\{\mathbf{r}_k\}_{k=1:K}$$

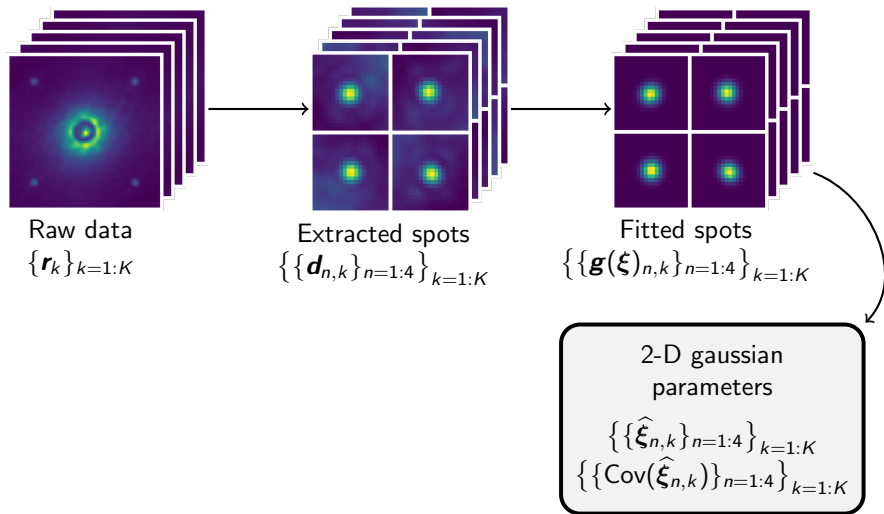
Extracting & fitting the satellite spots



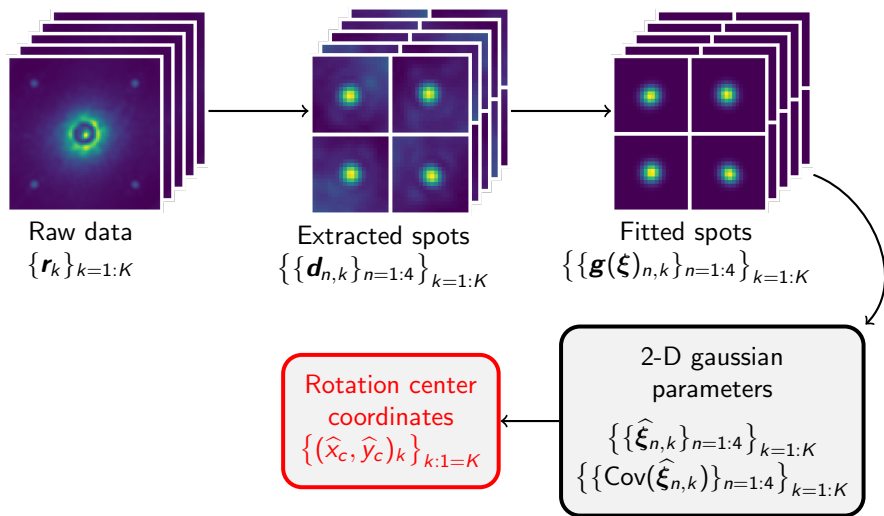
Extracting & fitting the satellite spots



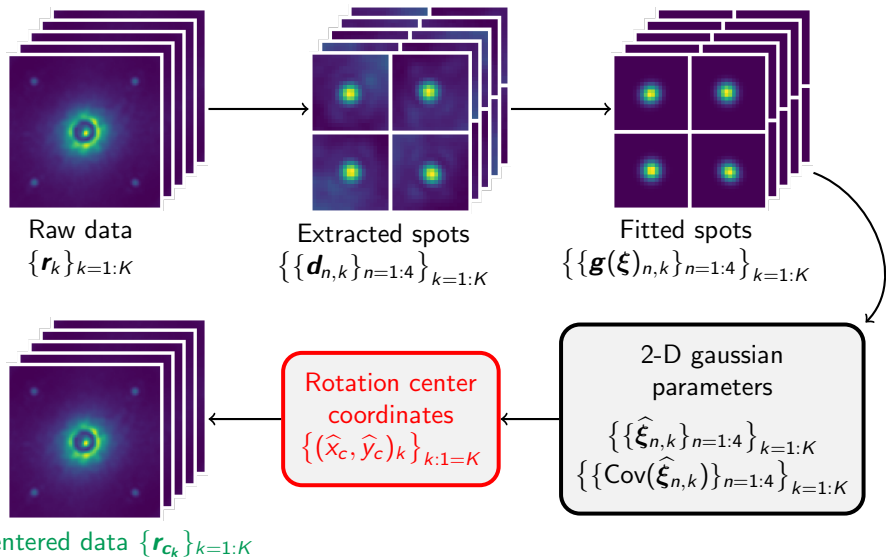
Extracting & fitting the satellite spots



Extracting & fitting the satellite spots



Extracting & fitting the satellite spots



Some details of the fitting procedure

- The satellite spots are extracted in 30 pixels-wide squares centered on their theoretical positions ($\simeq 14\lambda/D$ away from the theoretical rotation center)

Some details of the fitting procedure

- The satellite spots are extracted in 30 pixels-wide squares centered on their theoretical positions ($\simeq 14\lambda/D$ away from the theoretical rotation center)
- For IRDIS, if a satellite fit doesn't converge a median spatial filtering is applied to the extracted image and the fit is reprocessed on it

Some details of the fitting procedure

- The satellite spots are extracted in 30 pixels-wide squares centered on their theoretical positions ($\simeq 14\lambda/D$ away from the theoretical rotation center)
- For IRDIS, if a satellite fit doesn't converge a median spatial filtering is applied to the extracted image and the fit is reprocessed on it
- For IFS, the background contribution is stronger so the median spatial filter is always applied

Some details of the fitting procedure

- The satellite spots are extracted in 30 pixels-wide squares centered on their theoretical positions ($\simeq 14\lambda/D$ away from the theoretical rotation center)
- For IRDIS, if a satellite fit doesn't converge a median spatial filtering is applied to the extracted image and the fit is reprocessed on it
- For IFS, the background contribution is stronger so the median spatial filter is always applied
- A circular mask is systematically applied to the extracted images to get rid of the stellar leakages in the corners

Some details of the fitting procedure

- The satellite spots are extracted in 30 pixels-wide squares centered on their theoretical positions ($\simeq 14\lambda/D$ away from the theoretical rotation center)
- For IRDIS, if a satellite fit doesn't converge a median spatial filtering is applied to the extracted image and the fit is reprocessed on it
- For IFS, the background contribution is stronger so the median spatial filter is always applied
- A circular mask is systematically applied to the extracted images to get rid of the stellar leakages in the corners
- If one satellite spot is not fitted correctly, the center is estimated with the 3 remaining ones

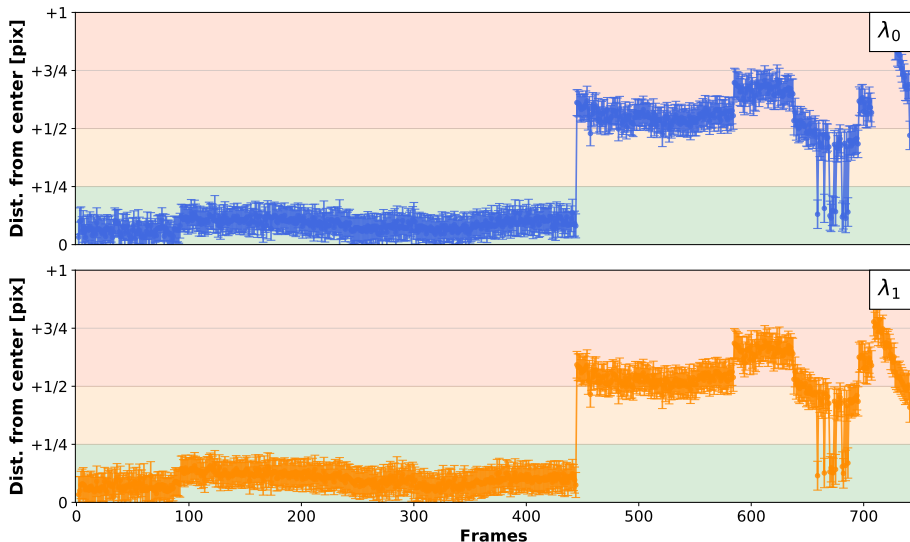
Some details of the fitting procedure

- The satellite spots are extracted in 30 pixels-wide squares centered on their theoretical positions ($\simeq 14\lambda/D$ away from the theoretical rotation center)
- For IRDIS, if a satellite fit doesn't converge a median spatial filtering is applied to the extracted image and the fit is reprocessed on it
- For IFS, the background contribution is stronger so the median spatial filter is always applied
- A circular mask is systematically applied to the extracted images to get rid of the stellar leakages in the corners
- If one satellite spot is not fitted correctly, the center is estimated with the 3 remaining ones
- Given the information on the spots pattern (\times or $+$) and the width of the band (BB or DB), some Gaussian parameters are fixed in the optimization procedure (θ , $\sigma_x = \sigma_y$)

Table of Contents

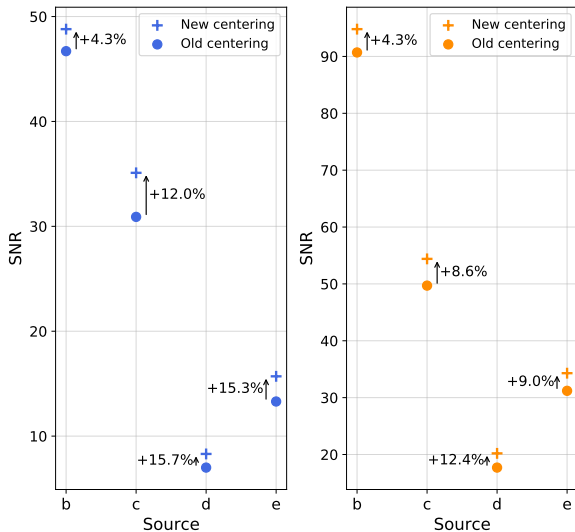
- ① Satellite spots
- ② Estimation of the rotation center
- ③ Application to SPHERE data
- ④ Conclusions & perspectives

Example of HR 8799, Dual-Band J2-J3 (2015-07-29)



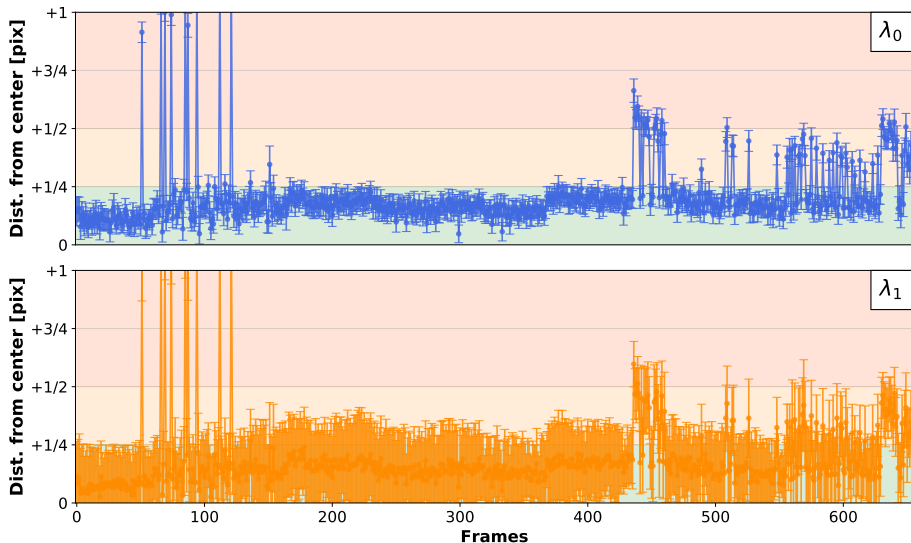
Dist. b/w the measured center of each frame and the theoretical rotation center

Example of HR 8799, Dual-Band J2-J3 (2015-07-29)



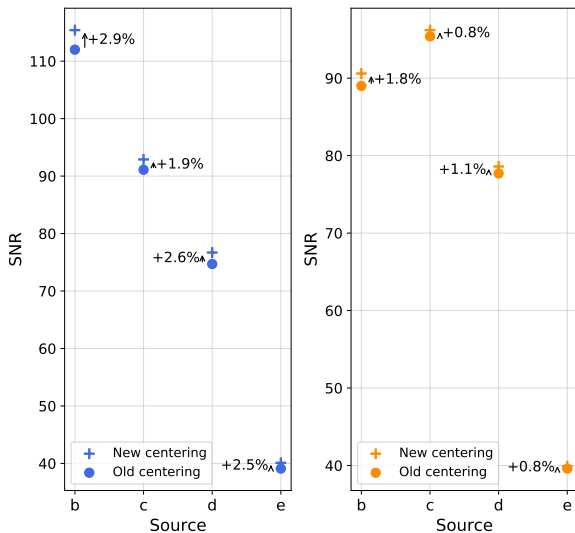
Comparison of PACO's SNR retrieved on the old dataset & the recentered one

Example of HR 8799, Dual-Band K1-K2 (2015-07-30)



Dist. b/w the measured center of each frame and the theoretical rotation center

Example of HR 8799, Dual-Band K1-K2 (2015-07-30)



Comparison of PACO's SNR retrieved on the old dataset & the recentered one

Table of Contents

- ① Satellite spots
- ② Estimation of the rotation center
- ③ Application to SPHERE data
- ④ Conclusions & perspectives

Fitting the satellite spots

- allows the accurate measurement of the rotation center
- enables the fine recentering of each individual frames of the cube
- impacts positively the SNR of off-axis point sources (for small shifts too)

Fitting the satellite spots

- allows the accurate measurement of the rotation center
- enables the fine recentering of each individual frames of the cube
- impacts positively the SNR of off-axis point sources (for small shifts too)

The fitting procedure gives access to the amplitude of the satellite spots

- could be used to calibrated the cube (real gain in SNR ?)
- but need to rethink the spatial filtering for IFS...

Fitting the satellite spots

- allows the accurate measurement of the rotation center
- enables the fine recentering of each individual frames of the cube
- impacts positively the SNR of off-axis point sources (for small shifts too)

The fitting procedure gives access to the amplitude of the satellite spots

- could be used to calibrated the cube (real gain in SNR ?)
- but need to rethink the spatial filtering for IFS...

It could be improved

- joint & simultaneous estimation of the 4 spots
- improve the robustness with a weighted least mean square optimization
- change the model (not perfectly gaussian), 2-D Moffat ? Bessel function ?
- the extracted ROIs have to cover the satellite spots for the method to work...